

A Uniform Expansion

$$\frac{\Delta d}{d} = \frac{\Delta t}{t_H}$$

d Distance to object now

Δd Change in distance to object during time Δt

t_H Expansion timescale

$t_H =$ age of the Universe if the expansion rate has always been constant

Today's best value : $t_H = 13.8 \pm 0.6$ Gyr

(Spergel et al., 2006, astro-ph/0603449)

$$\frac{\Delta d}{d} = \frac{\Delta t}{t_H}$$

A Universe right now has expansion timescale $t_H = 13.8$ Gyr. Assume the expansion rate has always been constant.

Galaxy A is twice as far away as Galaxy B. How will t_H as measured for Galaxy A compare to t_H as measured from for Galaxy B? (Assume that Galaxy A is no more than, say, a hundred million light-years away.)

A $t_H(A) = \frac{1}{4} t_H(B)$

B $t_H(A) = \frac{1}{2} t_H(B)$

C $t_H(A) = t_H(B)$

D $t_H(A) = 2 t_H(B)$

E $t_H(A) = 4 t_H(B)$

$$\frac{\Delta d}{d} = \frac{\Delta t}{t_H}$$

A Universe right now has expansion timescale $t_H = 13.8$ Gyr. Assume the expansion rate has always been constant. A certain galaxy is 100 Mpc away. How long has it been since that galaxy was 50 Mpc away?

A 3.45 Gyr

B 6.90 Gyr

C 13.8 Gyr

D 27.6 Gyr

E 55.2 Gyr

The expansion rate of the Universe is currently accelerating– that is, it is getting faster. It has been accelerating for something like the last 7 billion years.

Right now, we have $t_H = 13.6$ Gyr. (Remember that Gyr=gigayears=billion years.) 3.6 billion years ago, was the expansion timescale:

A <10.0 Gyr,

B Exactly 10.0 Gyr,

C between 10 Gyr and 13.6 Gyr,

D exactly 13.6 Gyr, or

E >13.6 Gyr?

C. I have no basis for this answer, but C is said to be most often right on multiple choice exams. I have no idea if this is correct, but I'm just going to follow my instincts.

$t_H=13.8$ Gyr right now. If the expansion rate of the Universe has been constant, what was t_H 3.8 Gyr ago?

- A $t_H < 10.0$ Gyr
- B $t_H = 10.0$ Gyr**
- C $10.0 \text{ Gyr} < t_H < 13.8 \text{ Gyr}$
- D $t_H = 13.8$ Gyr
- E $t_H > 13.8$ Gyr

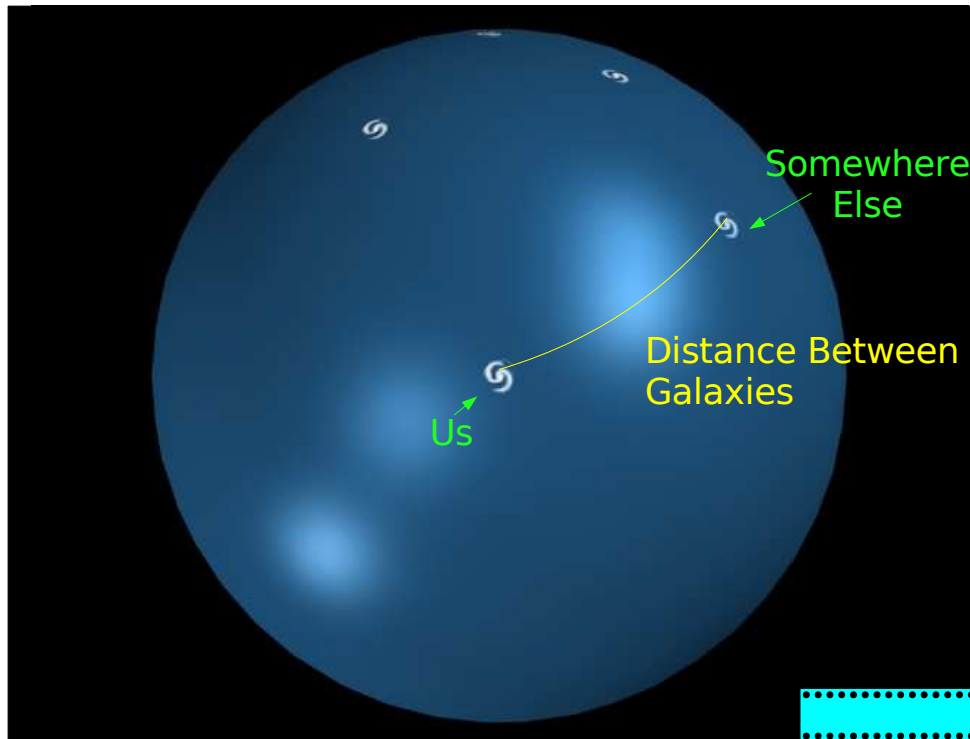
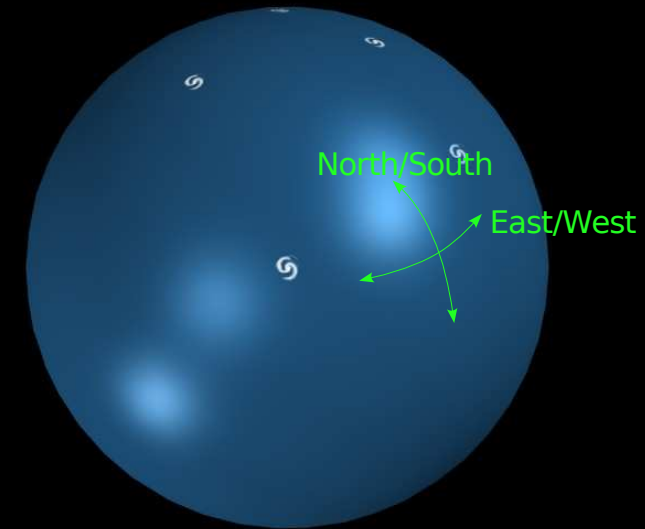
Does a *slower expansion rate* mean a *lower t_H* or a *higher t_H* ?

- A Lower
- B Higher**

If the expansion of the Universe is accelerating, what was t_H 3.8 Gyr ago?

- A $t_H < 10.0$ Gyr
- B $t_H = 10.0$ Gyr
- C $10.0 \text{ Gyr} < t_H < 13.8 \text{ Gyr}$**
- D $t_H = 13.8$ Gyr
- E $t_H > 13.8$ Gyr

A model 2-d closed Universe: the *surface* of a sphere



Redshift

$$z = \frac{\lambda_{\text{obs}} - \lambda_{\text{emit}}}{\lambda_{\text{emit}}} = \frac{\Delta\lambda}{\lambda}$$

This is the *definition* of observed redshift; whatever causes it, it is redshift.

$$1+z = \frac{\lambda_{\text{obs}}}{\lambda_{\text{emit}}}$$

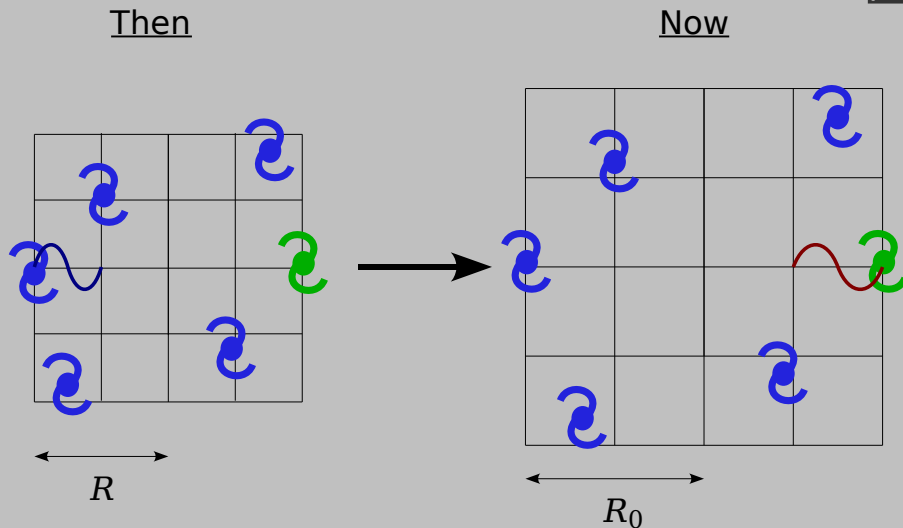
One possible source of redshift : the Doppler Shift

$$z = \frac{v}{c} \quad \text{for } v \ll c$$

$$1+z = \sqrt{\frac{1+\frac{v}{c}}{1-\frac{v}{c}}} \quad \text{for any } v, \text{ but I'll never make you use that}$$

Cosmological Redshift

As the Universe expands, the wavelength of light expands at exactly the same rate.



Cosmological Redshift

As the Universe expands, the wavelength of light expands at exactly the same rate.

$$\frac{\lambda_{\text{obs}}}{\lambda_{\text{emit}}} = 1 + z = \frac{R_0}{R}$$

R = "size of Universe" at emission (average distance between galaxies, or distance to a given galaxy)

R_0 = "size of Universe" now

λ_{emit} = emitted wavelength

λ_{obs} = observed (redshifted) wavelength

$$1 + z = \frac{d_0}{d}$$

$$z = \frac{d_0}{d} - 1 = \frac{d_0}{d} - \frac{d}{d}$$

$$z = \frac{d_0 - d}{d} = \frac{\Delta d}{d}$$

The cosmological redshift is a direct measurement of how much the Universe has expanded!!

$$\frac{\Delta d}{d} = \frac{\Delta t}{t_H}$$

$$z = \frac{\Delta t}{t_H}$$

We can also measure Δt – it's lookback time!

The Universe expands by $\Delta d = d_0 - d$ in time $\Delta t = d/c$

Δt = time for light to go distance d (time=distance/speed)

d = initial distance to galaxy – the distance the light goes

d_0 = distance now

$$\frac{\Delta d}{d} = \frac{\Delta t}{t_H}$$

$$z = \frac{\Delta t}{t_H}$$

Can Measure
(with a spectrum)

Can Measure
(dimmer std. candle = bigger d)

$$z = \frac{d}{c t_H}$$

Expanding Universe Equation
(only works for $z \ll 1$)

From previous measurements,
today we know this = 13.8 Gyr

Know